

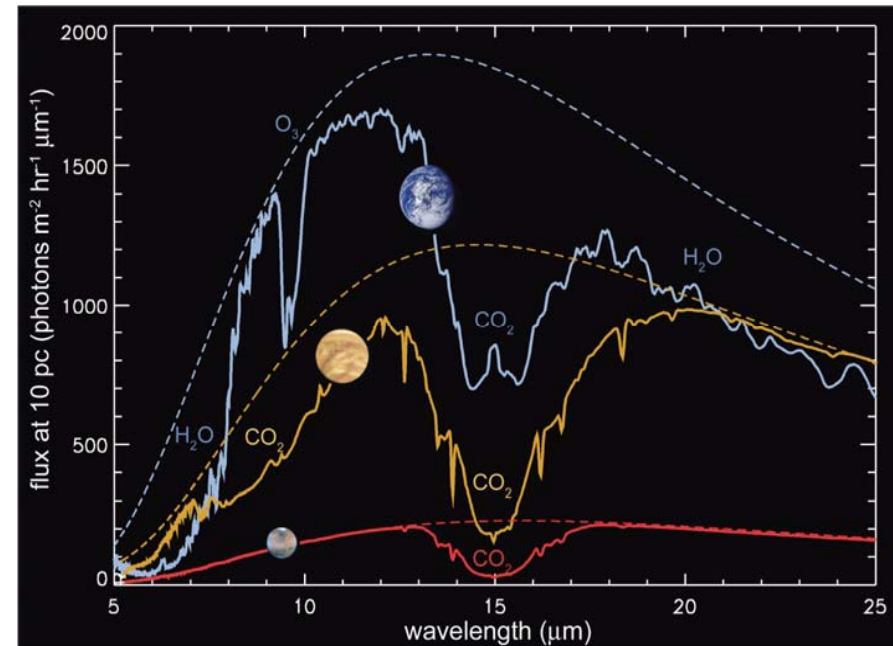
Emma X-Array Architecture for TPF-I / Darwin

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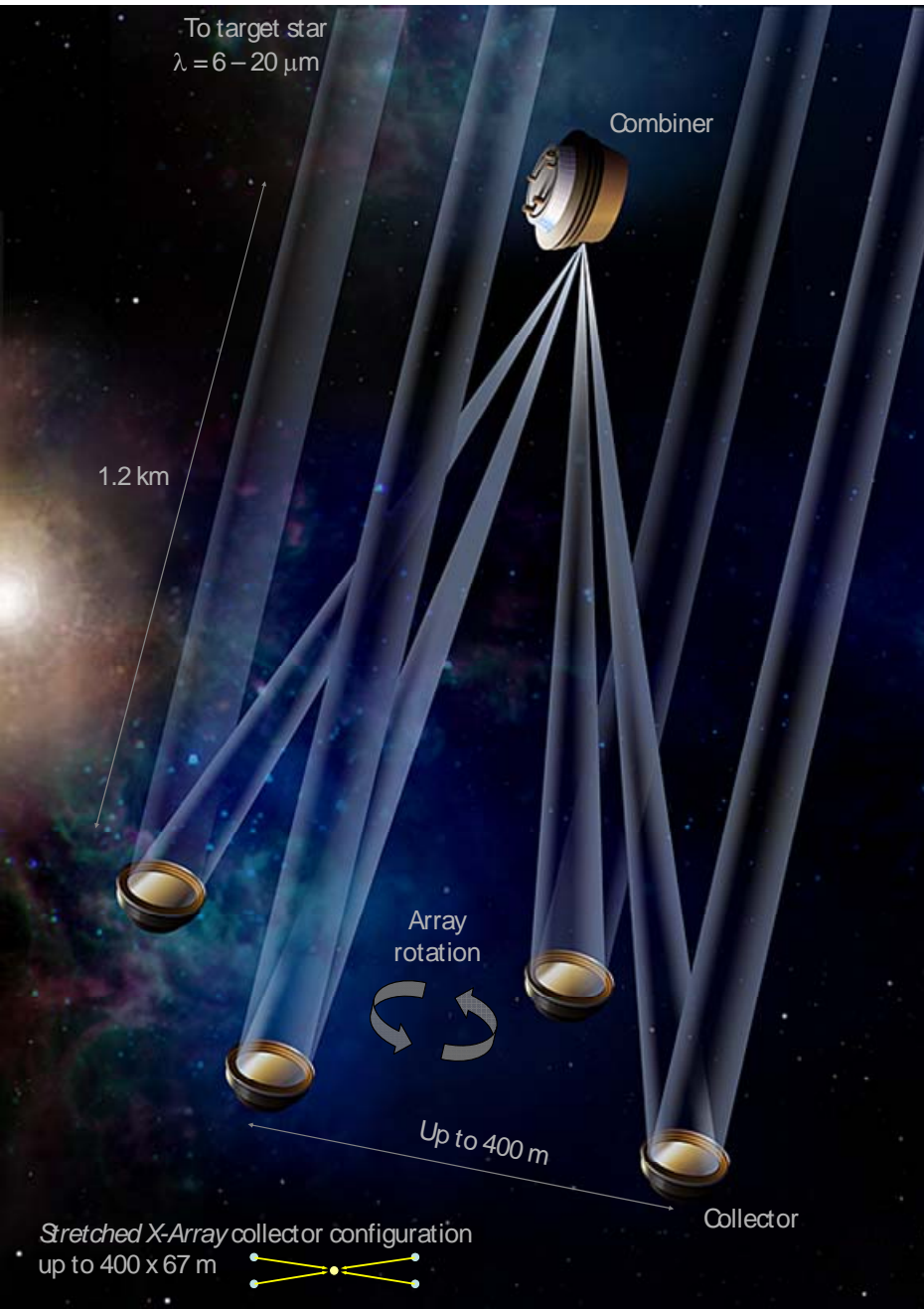
TPF-I / Darwin Key Features

- **Wavelengths:** Mid-Infrared 6 – 20 μm
- **Contrast:** Earth-Sun $\sim 10^7$ @ 10 μm
- **Biomarkers:** O_3 , CO_2 , CH_4 , H_2O
- **Technique:** Nulling Interferometry
- **Implementation:** Formation Flying

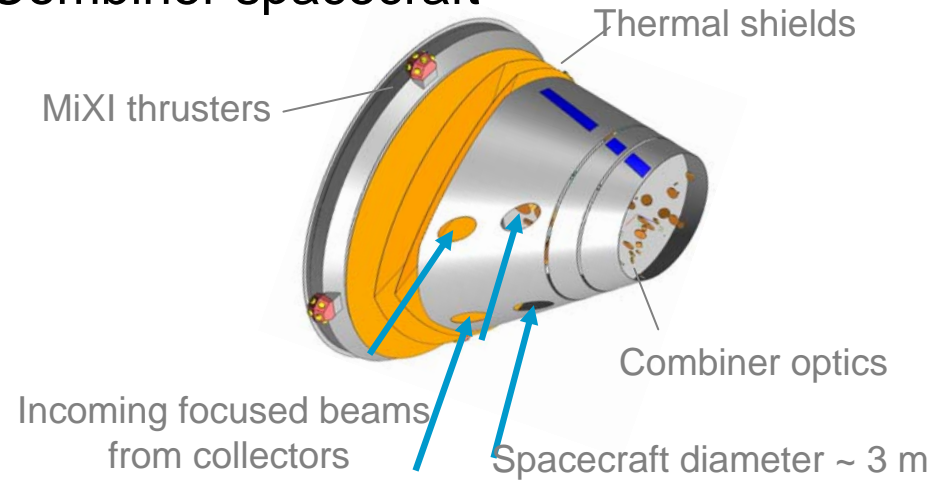


Selsis

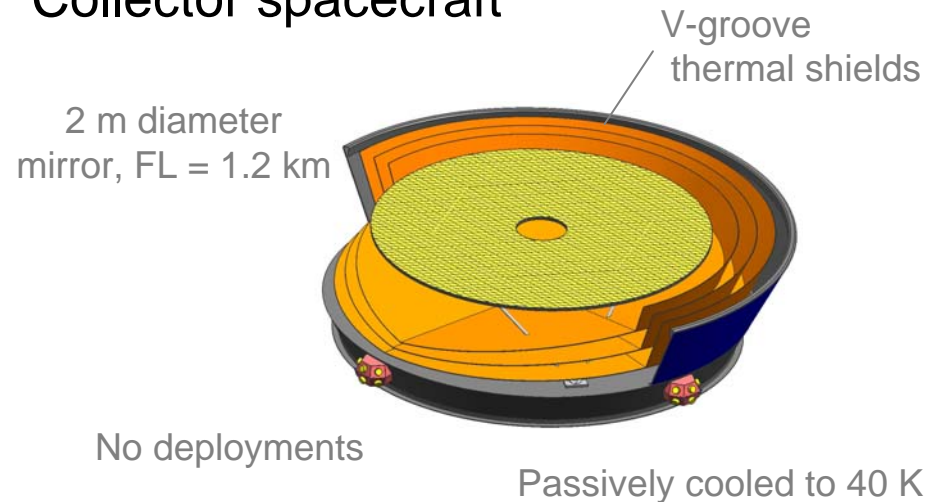
Emma X-Array Architecture



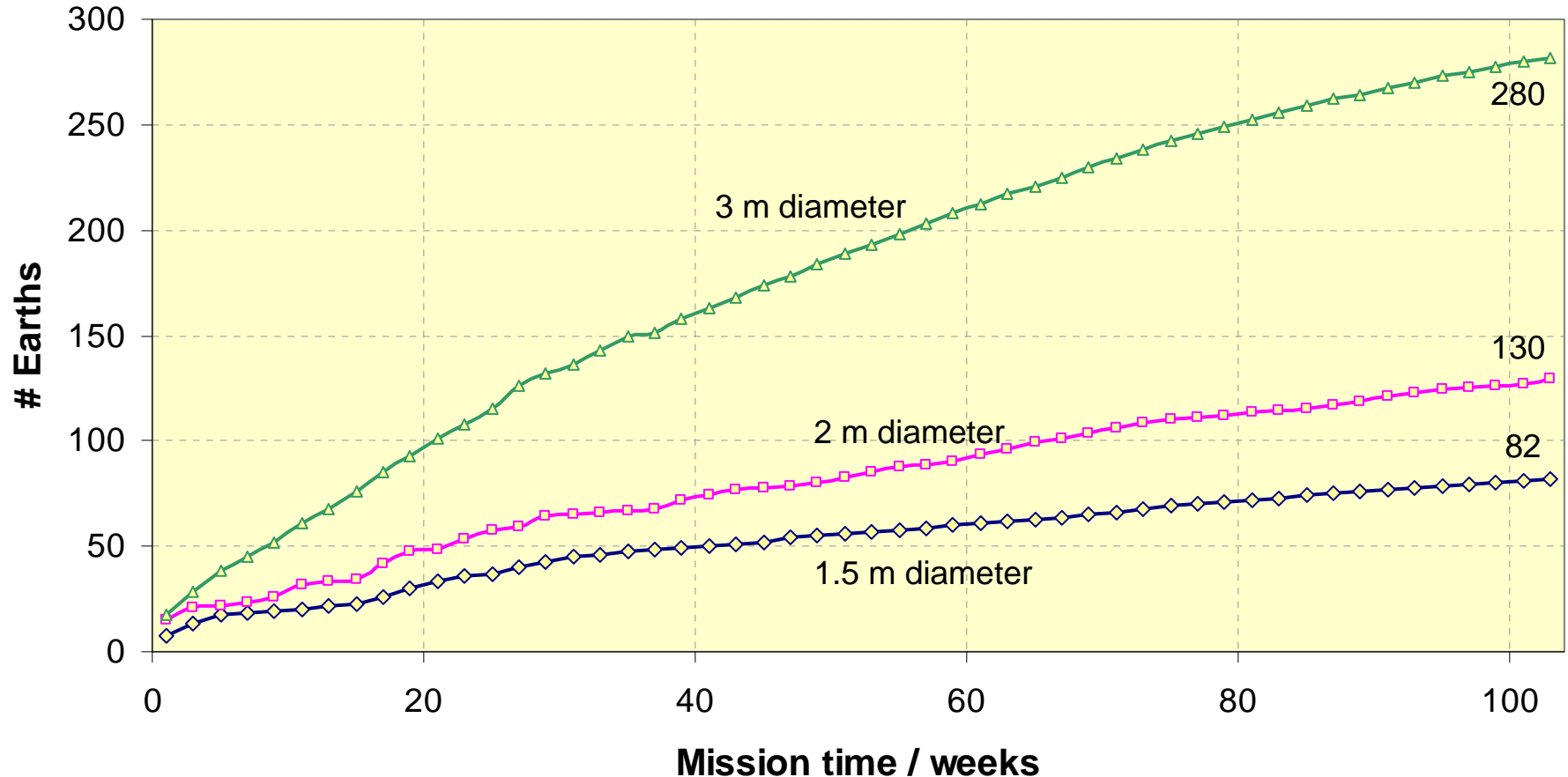
Combiner spacecraft



Collector spacecraft

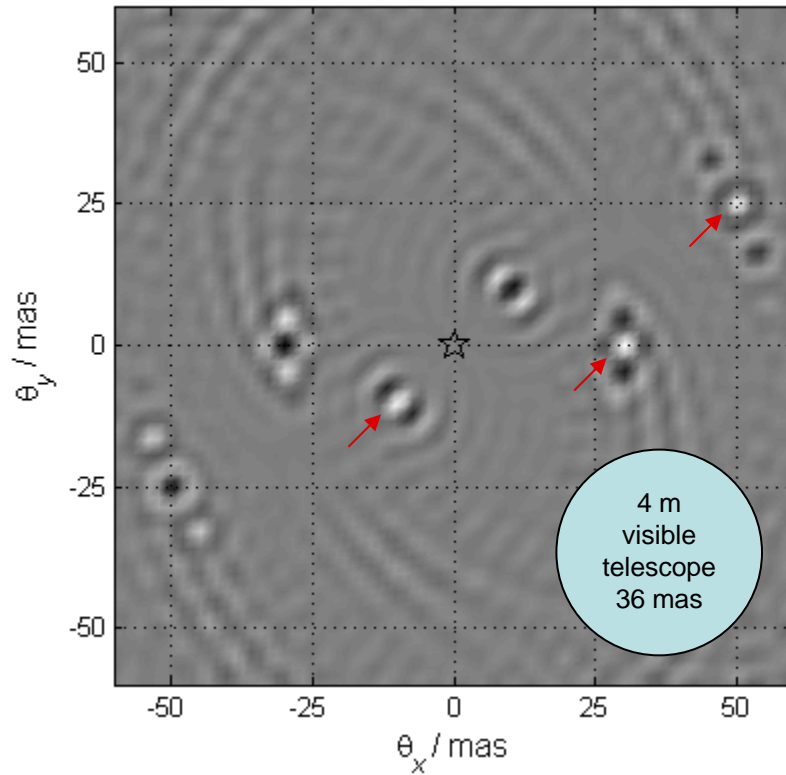


Earths found vs elapsed mission time

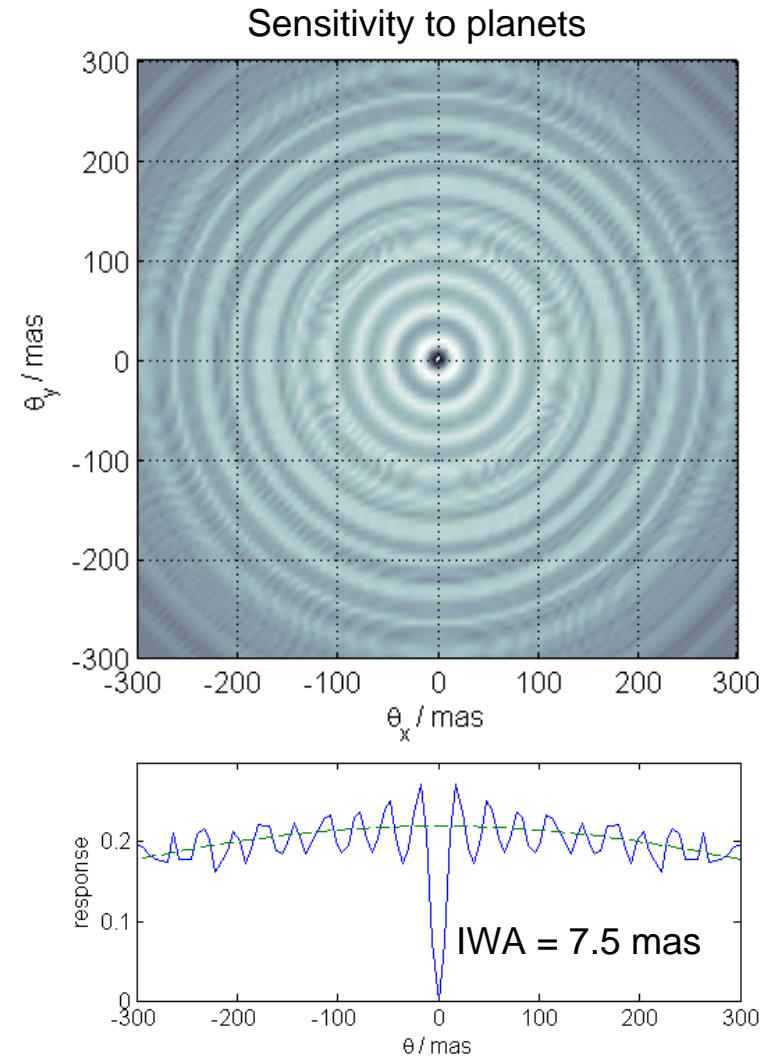


- 130 Earths after 2 years
- Assumes 1 Earth per star
- Only one visit per star in almost all cases

Imaging with 400 x 67 m array



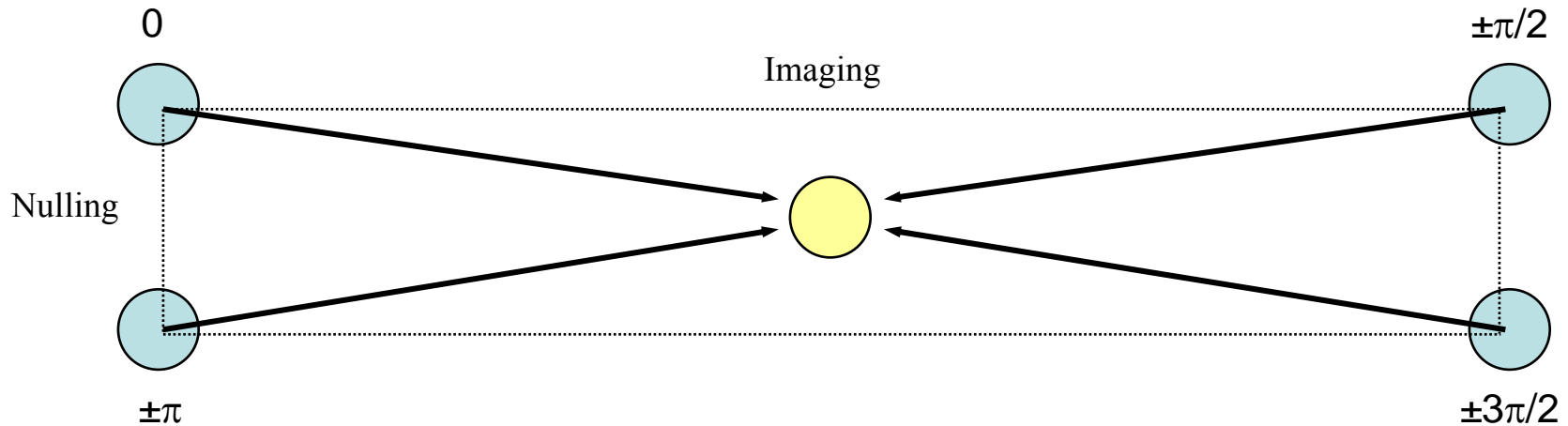
- Dirty map, no deconvolution
- 3 planets
- FWHM ~ 2.4 mas
- Separates multiple planets



- Inner Working Angle down to 7.5 mas

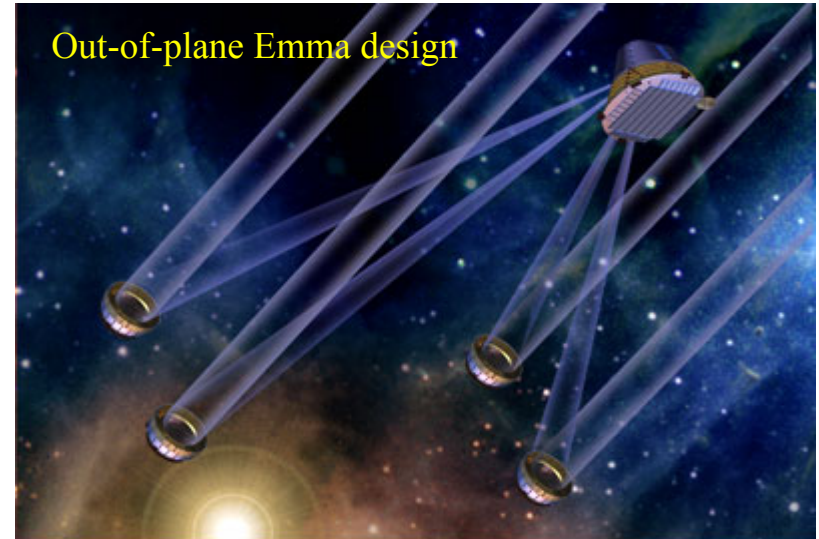
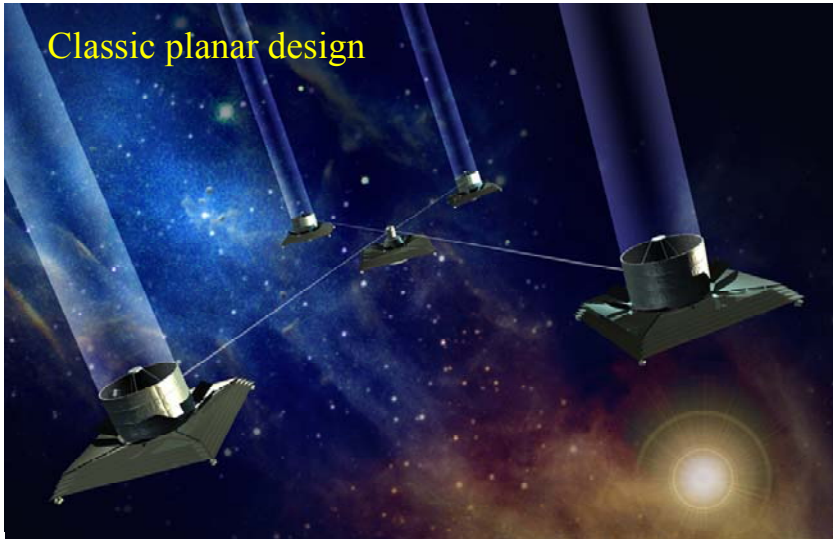
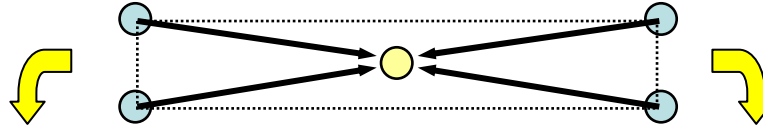
Back-up

Stretched X-Array



- $\sim 6:1$ aspect ratio
- Short nulling baselines, long imaging baselines
- Excellent angular resolution
- Can discriminate spectrally between planet signal and instrument 'instability noise'

Implementation



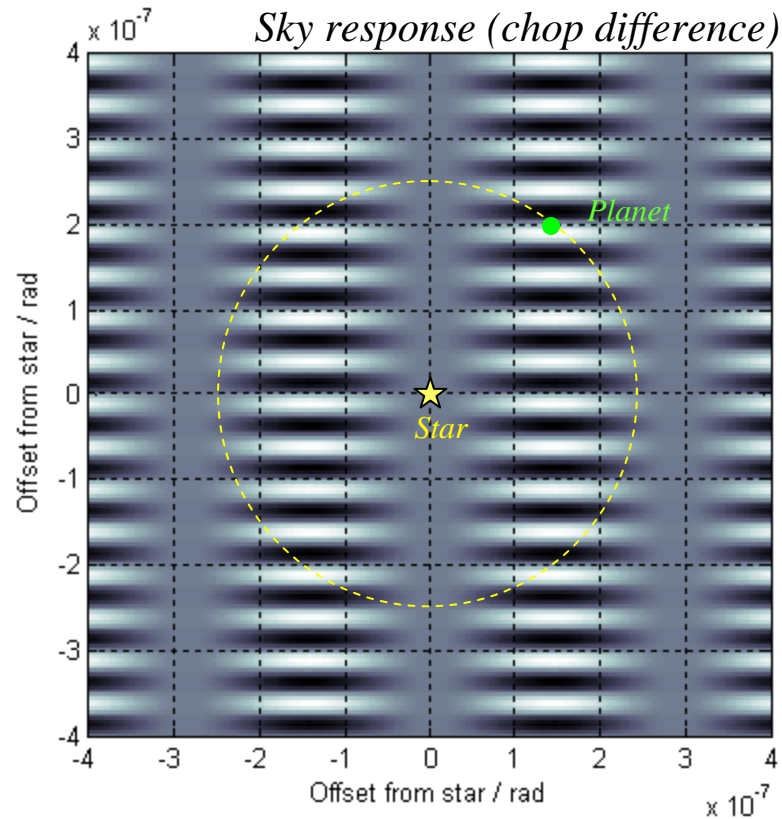
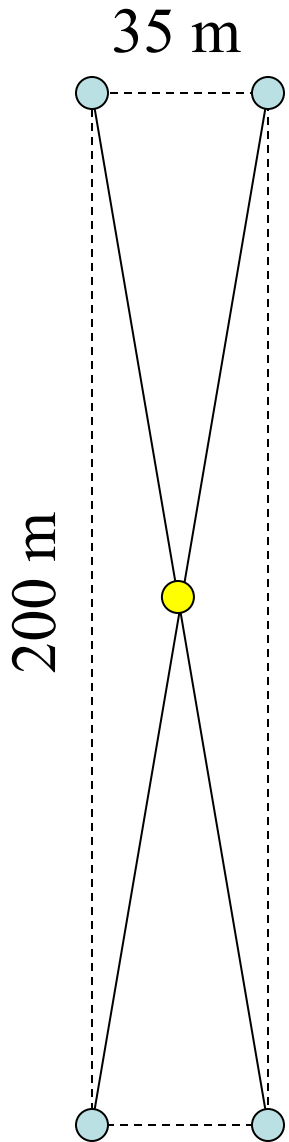
SPACECRAFT IMAGE BY T. HERBST (MPIA)

Performance discriminators

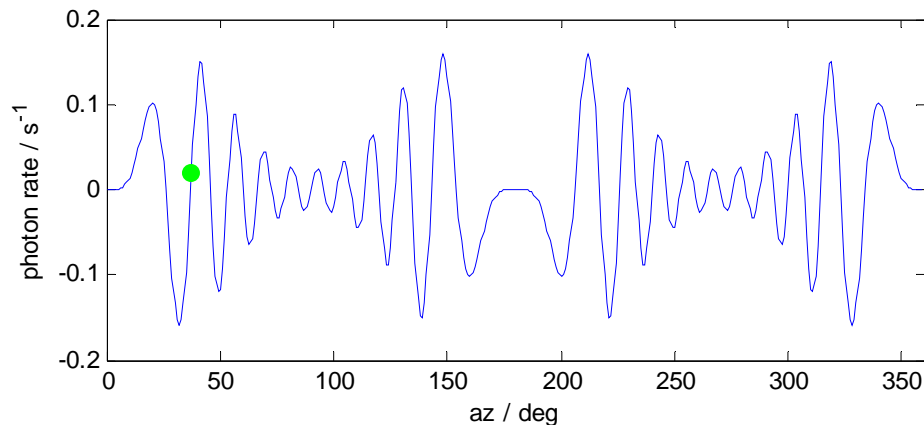
- Max. array size: 600 m
- Sky coverage: 71%
- Max. array size: 400 m
- Sky coverage: 99%

Response on the sky

X-Array 6:1

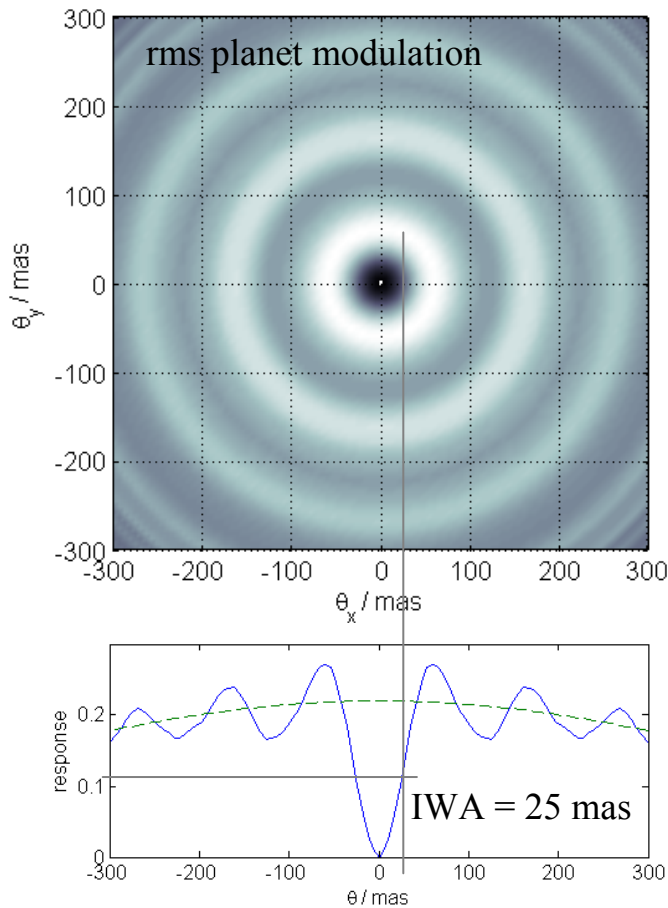


- $\lambda = 10 \mu\text{m}$
- Earth @ 50 mas
= 2.5×10^{-7} rad

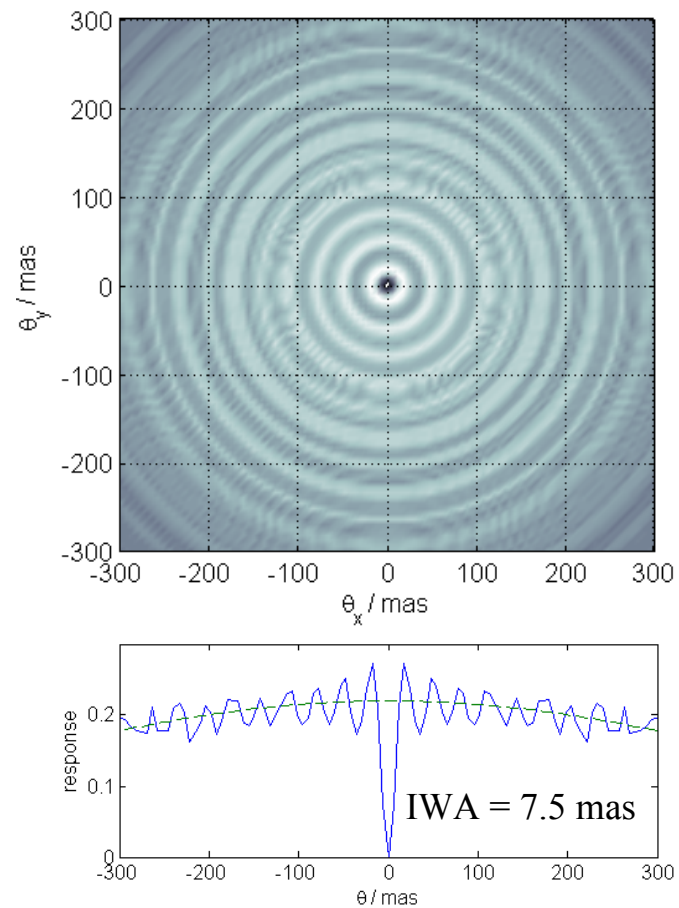


Inner Working Angle

120 x 20 m array



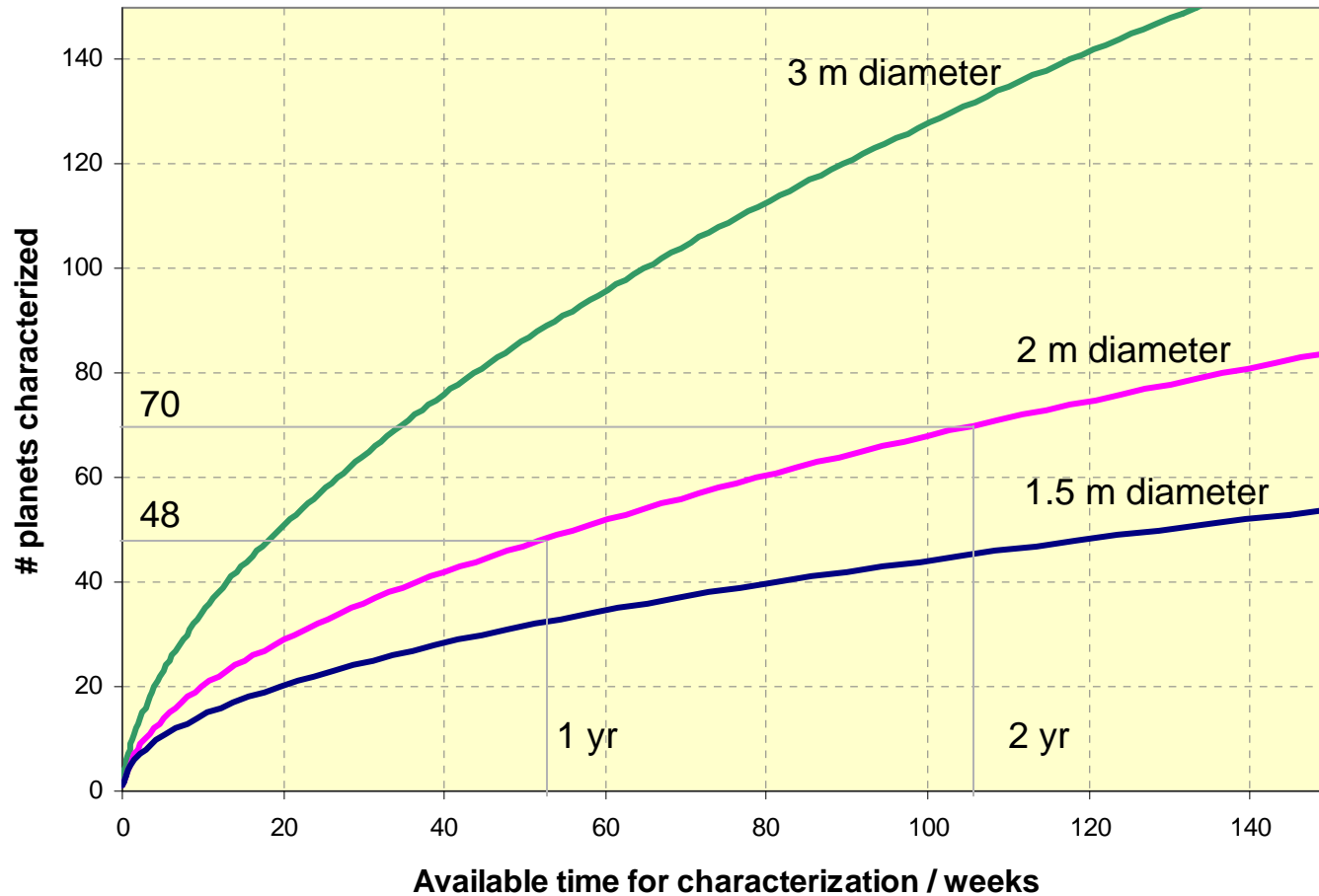
400 x 67 m array



Key observing parameters

Parameter	4-Telescope Chopped X-Array Emma Design
Collectors	Four spherical mirrors, diffraction limited at 2 μm operating at 50 K
Collector diameter	1.5, 2.0, 3.0 m
Array size	400 \times 67 m to 120 \times 20 m
Wavelength range	6–20 μm
Null depth	10 ⁻⁵ at 10 μm (not including stellar size leakage)
Exozodiacal emission	3 times our solar system, based on Kelsall model
Instrument throughput	10% (excludes ideal beam combiner losses)
Planet	Earth-sized 265 K black body
Retargeting time	6 hours
Integration efficiency	75%
SNR threshold	5

Spectroscopy



- SNR = 5 in 9.5 – 10 μm ozone channel
- 1 earth per star

Summary

- Emma X-Array design with 2 m collectors can detect ~ 130 Earths in 2 years
- Reduced array size is more than compensated by increased sky coverage
- Spectroscopy: ~ 50 nearest Earths in 1 year
- Excellent imaging with 2.4 mas angular resolution



SPACECRAFT IMAGE BY T. HERBST (MPIA)

Integration times for planet detection

